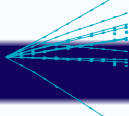


# Neutrinoless Double Beta Decay and COBRA

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20 May 2010



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- ① Neutrino Physics and the Neutrinoless Double Beta Decay
- ② Experimental Considerations
- ③ Experiments in Past, Presence and Future
- ④ The COBRA Experiment



# Neutrino Physics and the Neutrinoless Double Beta Decay

- What we know about Neutrinos
- What we don't know about Neutrinos
- Dirac and Majorana Particles
- The Neutrinoless Double Beta Decay ( $0\nu\beta\beta$ )



# What we know about Neutrinos I

- From  $Z^0$  decay:  $N_\nu = 2.980 \pm 0.025$
- Flavour eigenstates  $\neq$  mass eigenstates, they mix:

$$|\nu_\alpha\rangle = \sum_{i=1}^3 U_{\alpha i}^* |\nu_i\rangle$$

$U$  is called PMNS–Matrix (Pontecorvo, Maki, Nakagawa, Sakata).

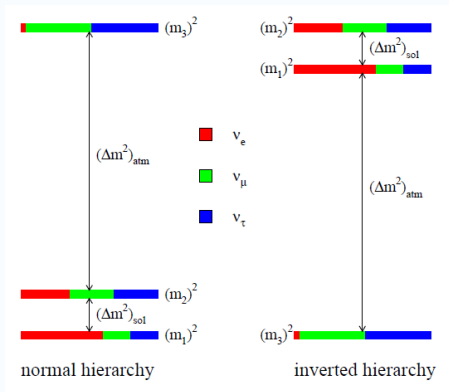
Parameters (values from Schwetz *et al.*, *New. J. Phys.* 10 (2008) 113011.):

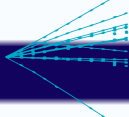
- $\sin^2(\theta_{12}) = 0.304_{-0.016}^{+0.022}$
- $\sin^2(\theta_{23}) = 0.50_{-0.06}^{+0.07}$
- $\sin^2(\theta_{13}) = 0.01_{-0.011}^{+0.016}$
- 1 CP–violating phase
- 2 additional phases if neutrinos are Majorana particles

# What we know about Neutrinos II

From oscillation experiments (*Schwetz et al., New. J. Phys. 10 (2008) 113011.*):

- $\Delta m_{21}^2 = \Delta m_{sol}^2 = 7.65_{-0.20}^{+0.23} \cdot 10^{-5} \text{ eV}^2$
- $|\Delta m_{31}^2| = |\Delta m_{atm}^2| = 2.40_{-0.11}^{+0.12} \cdot 10^{-3} \text{ eV}^2$



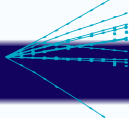


# What we don't know

Open questions:

- What is the nature of neutrinos (Dirac vs. Majorana)?
- Is the charge/parity (CP) symmetry broken?
- What are the precise values of neutrino masses and mixing?
- Are there sterile neutrinos? Is the standard picture right?

A wide experimental program is going to address these questions in the future.



# What we don't know

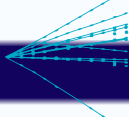
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A wide experimental program is going to address these questions in the future.

$0\nu\beta\beta$  experiments

- are the only known way to distinguish between Dirac and Majorana neutrinos
- can determine the absolute neutrino mass



# Dirac Particles

In quantum field theory:

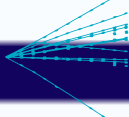
Spin-1/2 particles are described by four-component spinors which obey the Dirac equation.

The four independent components correspond to:

- Particles with Helicity  $\pm 1$
- Antiparticles with Helicity  $\pm 1$

All charged leptons ( $e$ ,  $\mu$ ,  $\tau$ ) have to be Dirac particles.





# Majorana Particles

Experimental fact:

Only left-handed neutrinos ( $H = -1$ ) and right-handed antineutrinos ( $H = +1$ ) are observed.

If neutrinos are Dirac particles:

- Two states are not realised in nature or are sterile

A two-component description should be sufficient (Weyl spinors).

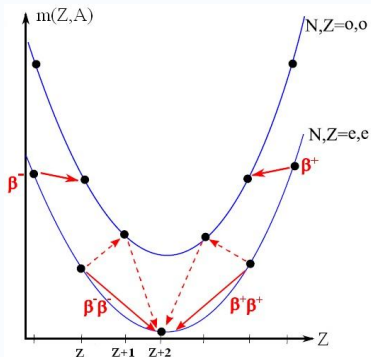
If neutrinos are Majorana particles:

- Particles = Antiparticles ( $\nu_L = \bar{\nu}_L, \nu_R = \bar{\nu}_R$ )
- Neutrinos don't carry a lepton number

# Double Beta Decay

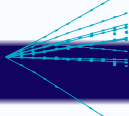
Goeppert–Mayer (1935):

- $(Z, A) \rightarrow (Z + 2, A) + 2e^- + 2\bar{\nu}_e$
- Simultaneous decay of two neutrons



Requirements:

- Nucleus with even  $A$ : Separation of  $m(Z, A)$  in 2 parabolas
- $m(Z, A) > m(Z + 2, A)$
- $m(Z, A) < m(Z + 1, A)$  (single  $\beta$  decay forbidden)



# Double Beta Decay

36  $2\beta^-$  isotopes are known today.

Furthermore possible:

- $\beta^+\beta^+$ :  $(Z, A) \rightarrow (Z - 2, A) + 2e^+ + 2\nu_e$  (6 isotopes)
- $\beta^+ + \text{EC}$ :  $(Z, A) + e^- \rightarrow (Z - 2, A) + e^+ + 2\nu_e$
- $\text{EC}/\text{EC}$ :  $(Z, A) + 2e^- \rightarrow (Z - 2, A) + 2\nu_e$

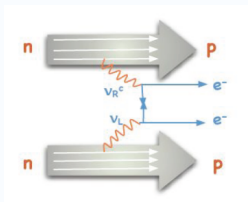
Definition *Q value*:

- Decay energy available for the leptons
- $Q = E_{e1} + E_{e2} + E_{\nu1} + E_{\nu2}$

$0\nu\beta\beta$  Decay

Furry (1939):

- $(Z, A) \rightarrow (Z + 2, A) + 2e^-$



Racah sequence:

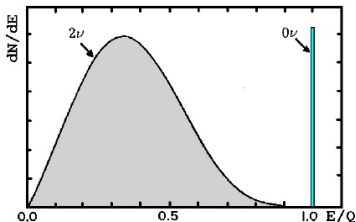
- $(Z, A) \rightarrow (Z + 1, A) + e^- + \bar{\nu}_{e,R}$   
 $(Z + 1, A) + \nu_{e,L} \rightarrow (Z + 2, A) + e^-$

Helicity flip is required: Suppression of  $0\nu\beta\beta$  by  $\sim 10^{-5}$

$0\nu\beta\beta$  Decay

Experimental confirmation:  
Peak at the Q value of the decay

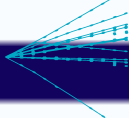
$$Q = E_{e1} + E_{e2}$$



Determination of neutrino mass from  $0\nu\beta\beta$ :

$$\left(T_{1/2}^{0\nu}\right)^{-1} = \overbrace{G^{0\nu}(Q, Z)}^{\text{phase space factor}} \cdot \underbrace{\left|M_{GT}^{0\nu} - M_F^{0\nu}\right|^2}_{\text{nuclear matrix element}} \cdot \overbrace{\frac{\langle m_{\nu_e} \rangle^2}{m_e^2}}^{\text{effective Majorana neutrino mass}}$$

$$\langle m_{\nu_e} \rangle = \left| \sum_i U_{ei}^2 m_i \right| \neq m_{\nu_e} = \sqrt{\sum_i |U_{ei}|^2 m_i^2}$$



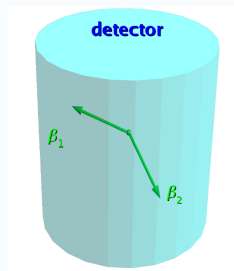
# Experimental Considerations

- General Concepts and Detection Limit
- Background Reduction
- The Favourite Isotopes

# General Detector Concepts

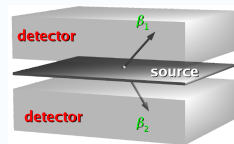
## Source = detector:

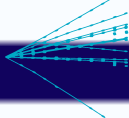
- Constraints on material selection
- + Big masses achievable
- + High energy resolution
- + For some isotopes reconstruction of event topology possible (LXe)



## Source $\neq$ detector:

- + Precise reconstruction of event topology
- + Analysis of different isotopes
- Big masses problematic



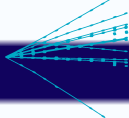


# Detection Limit

Without background:  $T_{1/2}^{limit} \propto a \cdot \epsilon \cdot M \cdot t$

- $a$  ... abundance
- $\epsilon$  ... detection efficiency
- $M$  ... mass
- $t$  ... measuring time



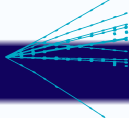


# Detection Limit

Without background:  $T_{1/2}^{limit} \propto a \cdot \epsilon \cdot M \cdot t$

With background:  $T_{1/2}^{limit} \propto a \cdot \epsilon \cdot \sqrt{\frac{M \cdot t}{\Delta E \cdot B}}$

- $a$  ... abundance
- $\epsilon$  ... detection efficiency
- $M$  ... mass
- $t$  ... measuring time
- $\Delta E$  ... energy resolution
- $B$  ... background event rate

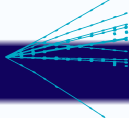


# Background

Benchmark for all future experiments:

Background  $< 10^{-3}$  counts/kg/keV/yr

For an experiment with  $M = 1$  ton: 1 count/yr at  $Q \pm 0.5$  keV



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Background  $< 10^{-3}$  counts /kg/keV/yr

For an experiment with  $M = 1$  ton: 1 count/yr at  $Q \pm 0.5$  keV

Main contributions to background:

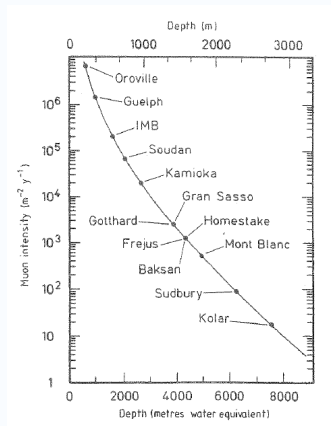
Cosmic rays/muons	→	underground laboratory
Cosmogenic radioisotopes	→	underground laboratory
Neutrons	→	shielding
Radioisotopes $^{238}\text{U}$ , $^{232}\text{Th}$ , $^{40}\text{K}$	→	shielding + clean environment
$2\nu\beta\beta$	→	high energy resolution

# Underground Laboratories

Rock-shielding decreases  $\mu$ -flux and hence cosmogenic radioisotopes

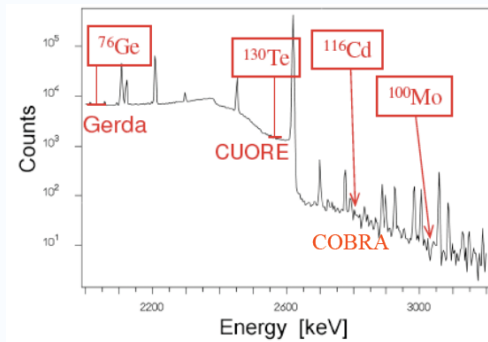
Laboratori Nazionali del Gran Sasso:

- World largest underground laboratories for particle physics
- e.g. OPERA, BOREXINO, DAMA, CRESST
- $0\nu\beta\beta$ : Heidelberg-Moscow, GERDA, COBRA, CUORE



# Radioisotopes

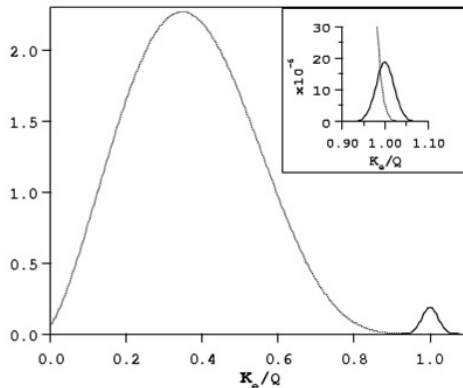
Highest  $\gamma$ -peak from  $^{232}\text{Th}$ : 2614 keV



Significant decrease of  $\gamma$ -background with  $Q > 2614$  keV

# Energy Resolution

Good energy resolution helps to separate  $0\nu\beta\beta$ -peak from  $2\nu\beta\beta$ -background





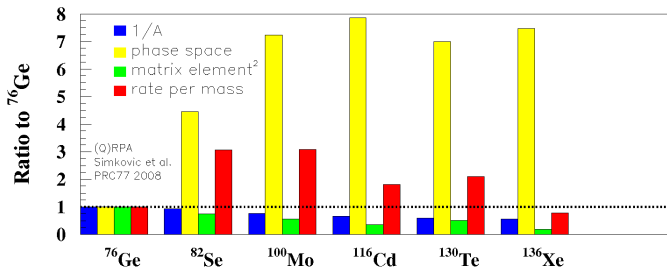
# Half-life

For a given  $\langle m_{\nu_e} \rangle$  the expected half-life can be calculated:

$$\left(T_{1/2}^{0\nu}\right)^{-1} = G^{0\nu}(Q, Z) \cdot \left|M_{GT}^{0\nu} - M_F^{0\nu}\right|^2 \cdot \frac{\langle m_{\nu_e} \rangle^2}{m_e^2}$$

- Phase space factor  $G^{0\nu}(Q, Z)$ 
  - precise calculation possible
- Nuclear matrix element  $\left|M_{GT}^{0\nu} - M_F^{0\nu}\right|$ 
  - calculations are not precise and model-dependent
  - $\rightarrow$  different isotopes have to be investigated

# Combination of Matrix Elements and Phase Space



Expected  $0\nu\beta\beta$  rates per mass vary with a factor  $\approx 4$ .



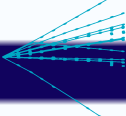
# Favoured Isotopes

Requirements for an experiment:

- High detection limit of the detector
- Low calculated half-life of the isotope

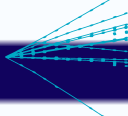
36  $2\beta^-$  isotopes are known today, but only 9 are considered for  $0\nu\beta\beta$  experiments:

		nat. abundance	Q [keV]
$^{48}\text{Ca}$	$\rightarrow$ $^{48}\text{Ti}$	0.19 %	4271
$^{150}\text{Nd}$	$\rightarrow$ $^{150}\text{Sm}$	5.6 %	3367
$^{100}\text{Mo}$	$\rightarrow$ $^{100}\text{Ru}$	9.6 %	3034
$^{82}\text{Se}$	$\rightarrow$ $^{82}\text{Kr}$	9.2 %	2995
$^{116}\text{Cd}$	$\rightarrow$ $^{116}\text{Sn}$	7.5 %	2809
$^{130}\text{Te}$	$\rightarrow$ $^{130}\text{Xe}$	33.8 %	2529
$^{136}\text{Xe}$	$\rightarrow$ $^{136}\text{Ba}$	8.9 %	2479
$^{124}\text{Sn}$	$\rightarrow$ $^{124}\text{Te}$	5.6 %	2288
$^{76}\text{Ge}$	$\rightarrow$ $^{76}\text{Se}$	7.8 %	2039



# Experiments in Past, Presence and Future

- The Heidelberg–Moscow Experiment
- Overview on Upcoming Experiments
- The GERDA Experiment



# Heidelberg–Moscow Overview

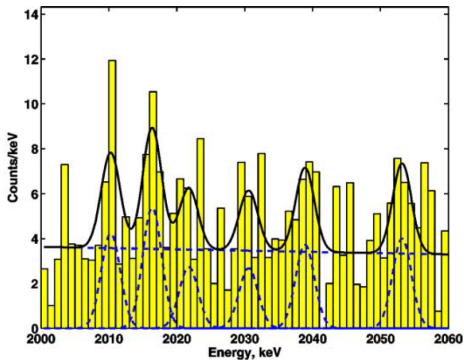
- Five p-doped HPGe semiconductor detectors
- 10.96 kg active mass
- $^{76}\text{Ge}$  enriched to 86%
  
- 1986: Proposal for this experiment by Prof. Klapdor-Kleingrothaus (Uni Heidelberg)
- 1990: Installation of first detector at LNGS (Central Italy)
- 1995: Completion of the experimental setup
- 2001: Exit of Kurchatov Institute
- 2003: End of data taking

# Setup Heidelberg–Moscow



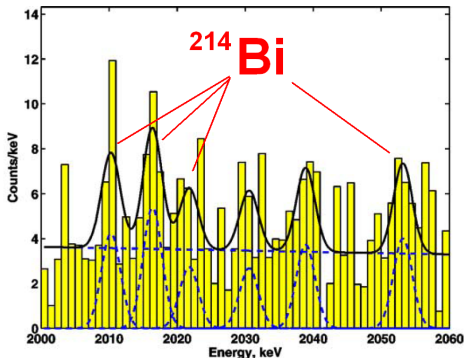
# Results $0\nu\beta\beta$ 2004

*H.V. Klapdor-Kleingrothaus et al, Phys.Lett.B586:198-212,2004*



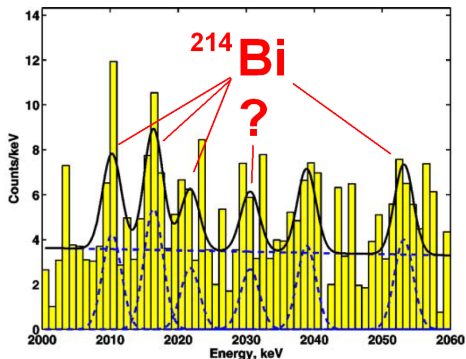
Results  $0\nu\beta\beta$  2004

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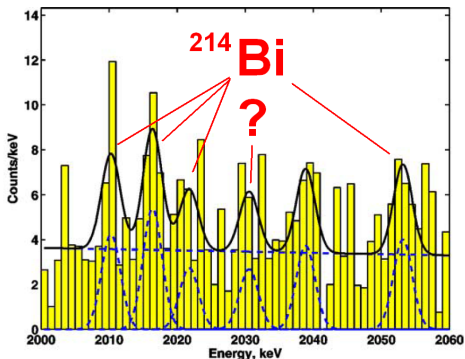
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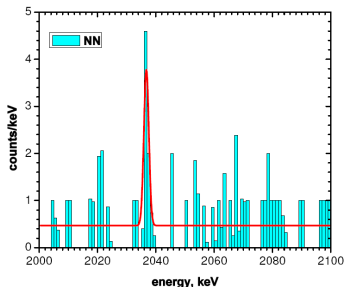
Claim of  $4.2\sigma$  evidence for a peak at 2038.07 keV

$$\rightarrow T_{1/2}^{0\nu} = 1.2 \cdot 10^{25} \text{ yr} \rightarrow \langle m_{\nu_e} \rangle = 0.44 \text{ eV}$$



# Results $0\nu\beta\beta$ 2006

*H.V. Klapdor-Kleingrothaus et al, Mod.Phys.Lett.A21:1547-1566,2006*



- New results with neural network and pulse shape analysis
- Claim of  $6.4\sigma$  evidence for a peak at  $\sim 2039$  keV
- Results are controversially discussed
- This claim has to be confirmed with other  $^{76}\text{Ge}$  experiments

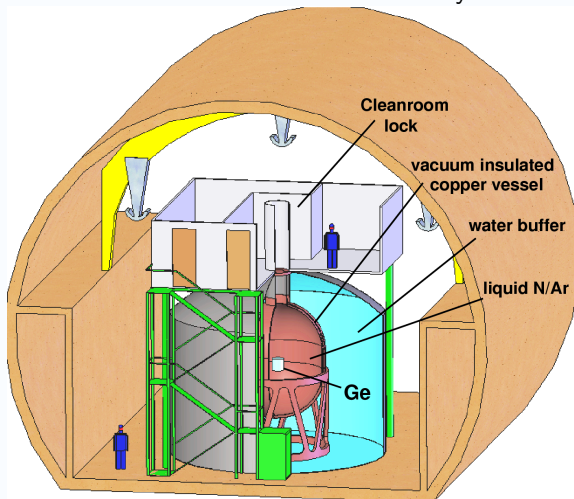


# Future Experiments

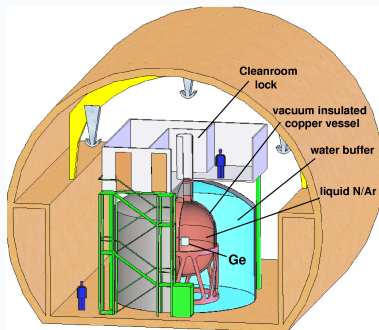
Name	Nucleus	Mass*	Method	Location	Time line
<i>Operational &amp; recently completed experiments</i>					
CUORICINO	Te-130	11 kg	bolometric	LNGS	2003-2008
NEMO-3	Mo-100/Se-82	6.9/0.9 kg	tracko-calo	LSM	until 2010
<i>Construction funding</i>					
CUORE	Te-130	200 kg	bolometric	LNGS	2012
EXO-200	Xe-136	160 kg	liquid TPC	WIPP	2009 (comiss.)
GERDA I/II	Ge-76	35 kg	ionization	LNGS	2009 (comiss.)
SNO+	Nd-150	56 kg	scintillation	SNOLab	2011
<i>Substantial R&amp;D funding / prototyping</i>					
CANDLES	Ca-48	0.35 kg	scintillation	Kamioka	2009
Majorana	Ge-76	26 kg	ionization	SUSL	2012
NEXT	Xe-136	80 kg	gas TPC	Canfranc	2013
SuperNEMO	Se-82 or Nd-150	100 kg	tracko-calo	LSM	2012 (first mod.)
<i>R&amp;D and/or conceptual design</i>					
CARVEL	Ca-48	tbd	scintillation	Solotvina	
COBRA	Cd-116, Te-130	tbd	ionization	LNGS	
DCBA	Nd-150	tbd	drift chamber	Kamioka	
EXO gas	Xe-136	tbd	gas TPC	SNOLab	
MOON	Mo-100	tbd	tracking	Oto	

## GERDA

## GERmanium Detector Array



## GERDA



- 2 m / 64 m<sup>3</sup> liquid Ar as inner shielding and for cooling
- 3 m / 650 m<sup>3</sup> high-purity water: Shielding against neutrons and  $\mu$ -veto (Water-Cherenkov-Detector)
- Clean room for preparation of detectors without contamination



# Phases

## Phase I

- 15 kg  $^{76}\text{Ge}$  detectors - background evaluation
- 18 kg  $^{76}\text{Ge}$  detectors - from Heidelberg–Moscow and IGEX
- Proof of Hd–M (2004) results within 1 year of data taking
- Expected: 6 events with 0.5 background events  $\rightarrow 5\sigma$



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## Phase II (simultaneous to Phase I)

- Development of new detectors
- Background reduction by 3 orders of magnitude
- Within 3 years of data taking:
  - Limit on  $T_{1/2}^{0\nu} : 2 \cdot 10^{26}$  yr
  - $0.09 \text{ eV} < \langle m_{\nu_e} \rangle < 0.29 \text{ eV}$



# Phases

## Phase I

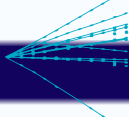
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## Phase III

- $\sim 1$  ton  $^{76}\text{Ge}$  (together with MAJORANA)
- $\langle m_{\nu_e} \rangle \sim 10 \text{ meV}$



# The COBRA Experiment

- The COBRA Collaboration
- Detector Concepts under Investigation
- CdZnTe Semiconductors in Liquid Scintillator





# The COBRA Collaboration



TU Dortmund

TU Dresden

Freiburger Material-  
forschungszentrum

Universität Hamburg

Universität Erlangen

Tschechische  
TU PragLaboratori Nazionali  
del Gran SassoWashington University  
at St. Louis

Universität Bratislava



Universität Jyväskylä



Universität La Plata



JINR Dubna

Kooperationen: Jagiellonen-Universität (Polen), Los Alamos  
Nat. Lab. (USA), University of Michigan (USA)

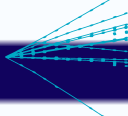


# Isotopes

## Cadmium–Zinc–Telluride $0\nu\beta\beta$ -neutrino double-Beta Research Apparatus

CdZnTe semiconductor detectors contain 9  $\beta\beta$ -isotopes:

Isotope	decay mode	Q (keV)	nat. abundance
$^{70}\text{Zn}$	$2\beta^-$	1001	0.62 %
$^{114}\text{Cd}$	$2\beta^-$	534	28.7 %
$^{116}\text{Cd}$	$2\beta^-$	2809	7.5 %
$^{128}\text{Te}$	$2\beta^-$	868	31.7 %
$^{130}\text{Te}$	$2\beta^-$	2529	33.8 %
$^{106}\text{Cd}$	$2\beta^+$	2771	1.21 %
$^{64}\text{Zn}$	$\beta^+/\text{EC}$	1096	48.6 %
$^{120}\text{Te}$	$\beta^+/\text{EC}$	1722	0.1 %
$^{108}\text{Cd}$	EC/EC	231	0.9 %



# Advantages of CdZnTe

- Source = detector (large mass)
- Semiconductor (good energy resolution, clean)
- Room temperature, no cryostat needed
- Modular design (coincidences)
- Tracking („Solid state TPC“)
- Industrial development of CdZnTe detectors
- Two isotopes at once
- $^{116}\text{Cd}$  above 2614 keV

# COBRA Concept

Aim:

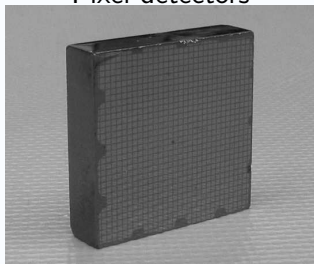
- Large-scale experiment with  $M = 420$  kg
- Main isotope:  $^{116}\text{Cd}$  enriched to  $\sim 90\%$
- Technical Design Report by end of 2012

Two detector concepts under investigation:

Coplanar grid detectors



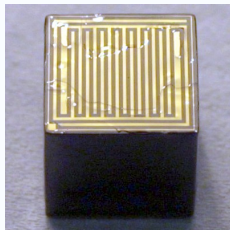
Pixel detectors



Energy measurement only

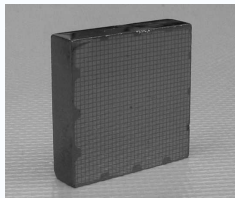
Energy measurement and tracking

# Coplanar Grid Detectors (CPG)



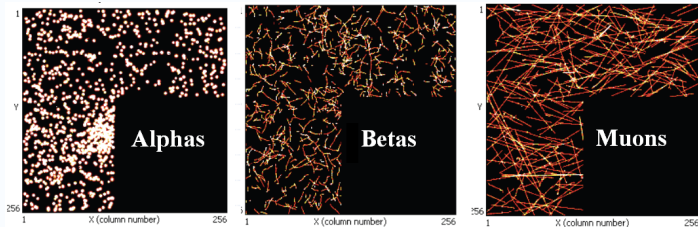
- Readout of two anodes to compensate signals from the holes (Trapping)
- CPGs are currently used in our test setup at LNGS
- Quite large volumes available (from  $1\text{ cm}^3$  to  $2\text{ cm}^3$ )
- Only few readout channels per mass
- Energy resolution around 2% FWHM at 2800 keV

# Pixel Detectors



- Particle identification: Background reduction by  $\approx 10^{-3}$

Real data with Si-Timepix (256x256 pixels):

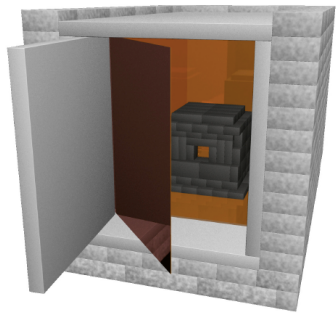
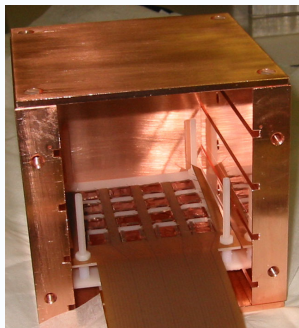


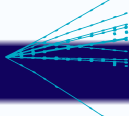
# Test Setup at LNGS



Upgrade summer 2010:

- 64 CPGs ( $4 \times 4 \times 4$ )
- Installation of Flash-ADCs
- New HV and LV

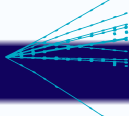




# Contribution of Uni Hamburg

- Shielding of large scale experiment
  - Monte Carlo studies
  - Design of alternative shielding concepts
  - Construction of prototypes and material selection
- Analysis of data from LNGS test setup
- Operation of CdZnTe in liquid scintillator





## CdZnTe in LSc: Motivation

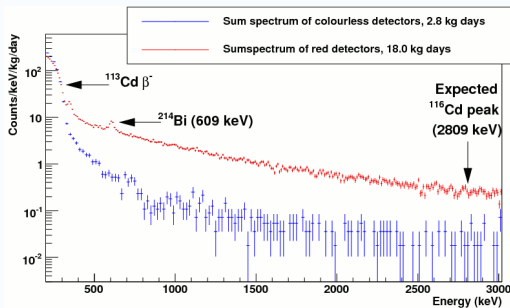
CPGs need passivation:

- to prevent CdZnTe from oxidation
- to avoid mechanical damage to the detector and the anodes

Passivations used so far:

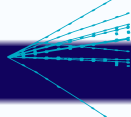
- Red lacquer (EI Detection & Imaging Systems)
- Transparent lacquer (EI Detection & Imaging Systems)
- Cyclotene (Dow Chemicals)

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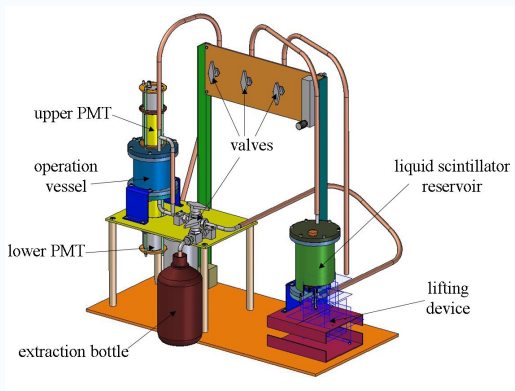


# Passivation with Liquid Scintillator

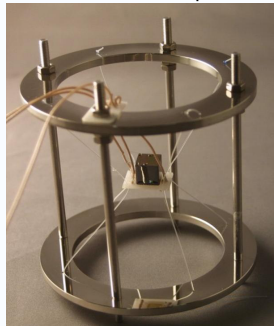
## Advantages of LSc:

- High impedance  $\rightarrow$  good isolator
- Offers possibility to control temperature stability
- Serves as active veto
- High purity with regard to radioactive nuclides  
 $\rightarrow$  e.g. BOREXINO or KamLAND

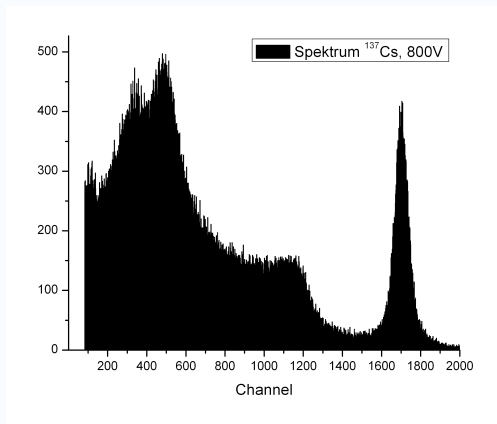
# Experimental Setup



1 cm<sup>3</sup> CdZnTe, not passivated:

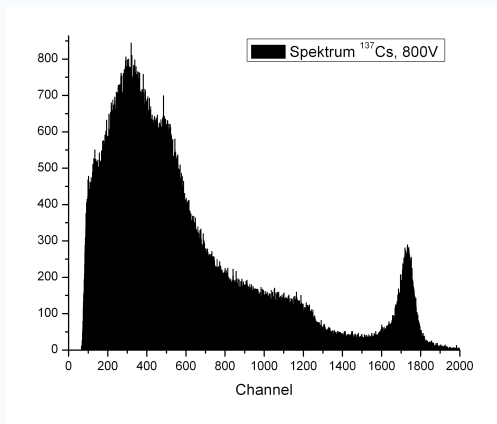


# $^{137}\text{Cs}$ in Nitrogen

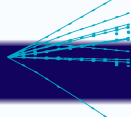


Energy resolution: 5.0% FWHM @ 662 keV

# $^{137}\text{Cs}$ in Liquid Scintillator

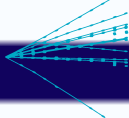


Energy resolution: 5.1% FWHM @ 662 keV



# First Results and Outlook

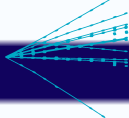
- CdZnTe detectors are operational in LSc with constant energy resolution
- Check for long term stability
- Investigation of different LSc mixtures
- Use LSc as active veto
- Upgrade of test setup to 8 detectors
- Operation of test setup in underground lab (Dresden or LNGS)



## Summary

- $0\nu\beta\beta$  experiments are the only known way to distinguish between Dirac and Majorana neutrinos
- Claim of  $0\nu\beta\beta$  discovery has to be confirmed
- Huge efforts are made for many different experiments
- COBRA can be the major and unique step beyond existing double beta experiments
- Operation of CdZnTe in liquid scintillator is an important contribution to the COBRA R&D programme





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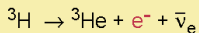
Thank you for your attention!

# The $\beta$ -decay and KATRIN

$\beta^-$  decay of the neutron

- $n \rightarrow p + e^- + \bar{\nu}_e$
- $(Z, A) \rightarrow (Z + 1, A) + e^- + \bar{\nu}_e$
- e.g. KATRIN:

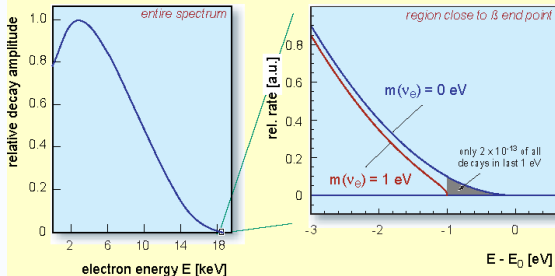
## tritium $\beta$ -decay and the neutrino rest mass



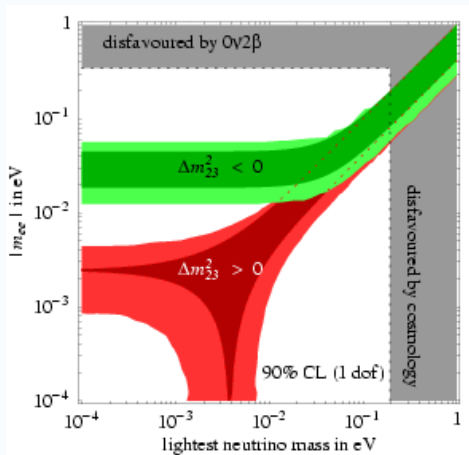
superallowed

half life :  $t_{1/2} = 12.32 \text{ a}$

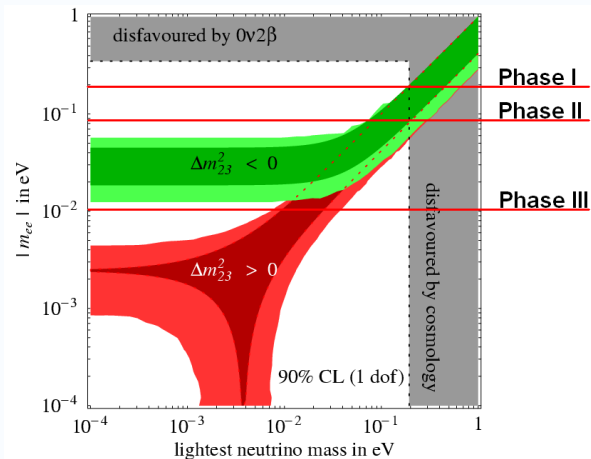
$\beta$  end point energy :  $E_0 = 18.57 \text{ keV}$



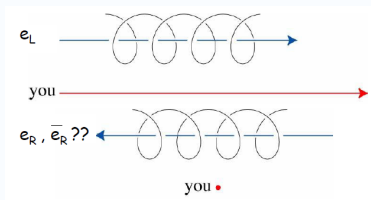
## Effective Neutrino Mass



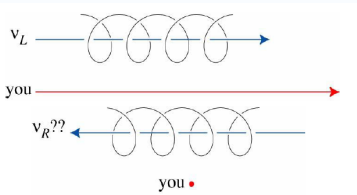
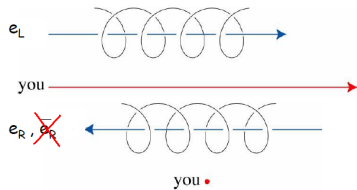
## Effective Neutrino Mass and GERDA



## Dirac vs. Majorana

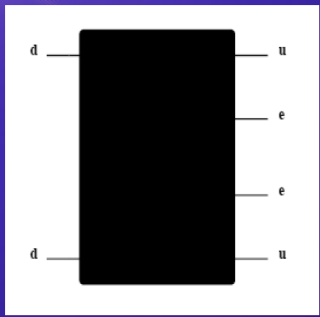


## Dirac vs. Majorana



## Other possible Mechanism

Further processes with  $\Delta L = 2$  that could contribute to  $0\nu\beta\beta$ :

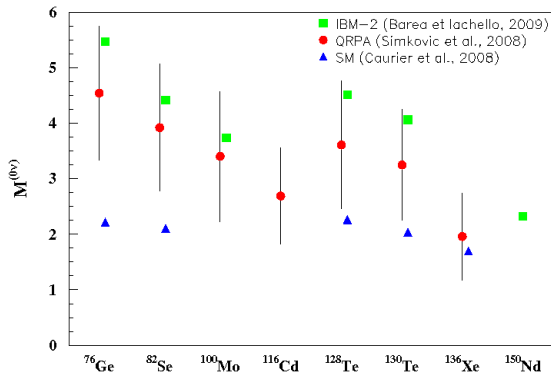


- $R_p$  violating SUSY
- V+A interactions
- Leptoquarks
- Double charged Higgs bosons
- Compositeness
- Heavy Majorana neutrino exchange
- Light Majorana neutrino exchange
- ...

Schechter and Valle, 1982:

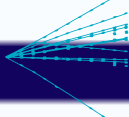
Observation of  $0\nu\beta\beta$  decay  $\equiv$  neutrinos are Majorana particles!

# Matrix Elements



Exact values of  $|M_{GT}^{0\nu} - M_F^{0\nu}|^2$  still not known.





# Natural Abundance

Enrichment of isotopes is very expensive

$^{48}\text{Ca}$	$\rightarrow$	$^{48}\text{Ti}$	0.2%
$^{76}\text{Ge}$	$\rightarrow$	$^{76}\text{Se}$	7.8%
$^{82}\text{Se}$	$\rightarrow$	$^{82}\text{Kr}$	9.2%
$^{100}\text{Mo}$	$\rightarrow$	$^{100}\text{Ru}$	9.6%
$^{116}\text{Cd}$	$\rightarrow$	$^{116}\text{Sn}$	7.5%
$^{124}\text{Sn}$	$\rightarrow$	$^{124}\text{Te}$	5.6%
$^{130}\text{Te}$	$\rightarrow$	$^{130}\text{Xe}$	34.5%
$^{136}\text{Xe}$	$\rightarrow$	$^{136}\text{Ba}$	8.9%
$^{150}\text{Nd}$	$\rightarrow$	$^{150}\text{Sm}$	5.6%