



### The OPERA Experiment Concluding the Neutrino Oscillation Analysis

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bmb+f - Förderschwerpunkt OPERA

Großgeräte der physikalischen Grundlagenforschung





### Neutrino oscillation in disappearance mode:

- First observation: SuperKamiokande, MACRO...
- Further studies: SNO, K2K, MINOS...

### Neutrino oscillation in appearance mode:

Observation needed to establish the picture of neutrino oscillations

### Solar scale:

•  $\nu_e \rightarrow \nu_\mu$ : Below threshold for  $\mu$  production

### Atmospheric scale:

- $u_{\mu} \rightarrow \nu_{\tau}$ :  $u_{\mu}$  from cosmic rays (SK, IceCube, ORCA)
- $\nu_{\mu} \rightarrow \nu_{\tau}$ :  $\nu_{\mu}$  from long-baseline beams OPERA: Event-by-event  $\tau$  lepton identification
- $\nu_{\mu} \rightarrow \nu_{e}$ : Sub-leading (T2K, OPERA)





## The **OPERA** Experiment

The OPERA experiment in the CERN to Gran Sasso neutrino beam, JINST **4** (2009) P04018





### OPERA: Oscillation Project with Emulsion Tracking Apparatus

- Appearance search: Direct observation of  $\nu_{\mu} \rightarrow \nu_{\tau}$  oscillations detection of  $\tau$  production & decay
- Characteristic 'kink' topology:



- $\nu$  beam: High-intensity & high-energy long-baseline  $\nu_{\mu}$  beam
- **Detector:** Large target mass, high precision  $\mathcal{O}(\mu m)$
- Location: Laboratori Nazionali del Gran Sasso (LNGS) 1400 m rock coverage, 3800 m w.e.

### The CNGS $\nu_{\mu}$ Beam



### CNGS: CERN Neutrinos to Gran Sasso (2008 - 2012)



p.o.t. (total)	$17.97  imes 10^{19}$	
$\nu$ interactions	19505	



45 50 E (GeV)

v<sub>n</sub> fluence

10 15 20 25 30 35 40

100

Hamburg University

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### The OPERA Detector





### Hybrid detector (ED & ECC):

- 2 identical Super Modules (SM) + VETO system
- Spectrometer: RPC & XPC, PT
- Target Area: TT, ECC bricks

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### The OPERA Detector





### Emulsion Cloud Chamber (ECC) bricks:

- 57  $\times$  2 AgBr nuclear emulsions on plastic bases, interleaved with 56 lead plates ( $\sim$  10  $X_0)$
- Total:  $~\sim 150\,000\,\times 8.3\,{\rm kg}~~\sim 1.25\,{\rm kt}$  total target mass
- Spatial / angular resolution:  $\sim 1 \,\mu m$  /  $\sim 2 \,mrad$

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### Changeable Sheets (CS):

• 2 extra nuclear emulsion sheets per brick

### Target Tracker (TT) detectors:

• Plastic scintillator strips (horizontal & vertical), 31 walls per SM



### The OPERA Detector





### Magnetic Spectrometer:

- Downstream of each target area
- Magnets: Iron core dipole, 1.55 T
- RPC, XPC: Resistive plate chambers
- Precision Tracker (PT):  $\sim 10\,000$  drift tubes

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### **Event Reconstruction**





### ED event reconstruction:

- Time resolution:  $\mathcal{O}(ns)$
- $\mu$  identification, charge & momentum measurement
- Hadronic shower energy reconstruction
- $\nu$  interaction brick localisation
- > Trigger: ECC event reconstruction



### **Event Reconstruction**





### ECC event reconstruction:

- Spatial resolution:  $\mathcal{O}(\mu m)$
- 3D track segment & track reconstruction
- $\nu$  interaction vertex localisation
- $\tau$  decay search procedure:
- ▷ kink angle / IP measurement, parent / daughter search...
- Momentum measurement via MCS





# **Oscillation Analysis:** $\nu_{\mu} \rightarrow \nu_{\tau}$

Discovery of  $\tau$  Neutrino Appearance in the CNGS Neutrino Beam with the OPERA experiment, Physical Review Letters **115**, 121802 (2015)



### The 1st $\nu_{\tau}$ Candidate Event:



- 1ry vertex: 7 tracks
- au candidate: 1-prong decay after  $(1335 \pm 35) \, \mu \mathrm{m}$
- $\triangleright$  Decay channel:  $\tau \rightarrow h$



### The Following 4 $\nu_{\tau}$ Candidates





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#### 28.03.2017 The OPERA Experiment

### $\nu_{\mu} \rightarrow \nu_{\tau}$ Oscillation Analysis



### Previous data sample (2008 – 2012): 5408 DS events

- 1st & 2nd most probable bricks
- All 0 $\mu$  events & 1 $\mu$  events with  $p_{\mu} < 15\,{\rm GeV/c}$

$\tau$ decay	Signal	Total BG	Data
channel	(exp.)	(exp.)	(obs.)
	$\Delta m_{23}^2 = 2.44 \mathrm{meV}^2$		
	$\sin^2 2 heta_{23} = 1$		
au  ightarrow 1 h	$0.52\pm0.10$	$\textbf{0.04} \pm \textbf{0.01}$	3
au  ightarrow 3h	$\textbf{0.73} \pm \textbf{0.14}$	$0.17\pm0.03$	1
$\tau \to \mu$	$0.61\pm0.12$	$0.004\pm0.001$	1
$\tau \to \mathbf{e}$	$0.78\pm0.16$	$\textbf{0.03} \pm \textbf{0.01}$	0
Total	$2.64\pm0.53$	$0.25\pm0.05$	5

### Discovery of $\nu_{\tau}$ appearance:

- p-value:  $1.10 \times 10^{-7}$  (Fisher) /  $1.07 \times 10^{-7}$  (profile likelihood)
- $\triangleright$  No-oscillation hypothesis excluded @  $5.1\sigma$



### $\nu_{\mu} \rightarrow \nu_{\tau}$ Oscillation Analysis



### Final data sample (2008 – 2012): 5603 DS events

- 1st & 2nd most probable bricks
- All 0 $\mu$  events & 1 $\mu$  events with  $p_{\mu} < 15\,{
  m GeV/c}$
- Increased statistics: +195 DS events

### Minimum bias analysis:

- Loosened kinematical cuts
- Boosted Decision Trees (kin. & topol. variables)
- Improved signal-to-noise ratio
- Increased statistics:

### $\triangleright$ 5 new $\nu_{\tau}$ candidates





### Measurement of $\Delta m_{23}^2$ in appearance mode:

- $N_{\nu_{\tau}} \propto (\Delta m_{23}^2)^2 L^2 \int \Phi(E) \epsilon(E) \frac{\sigma(E)}{E^2} dE$
- $\Delta m_{23}^2 = 2.95 \times 10^{-3} \,\mathrm{eV}^2$  ([1.98 3.95]  $\times 10^{-3} \,\mathrm{eV}^2$ ) PRELIM. for  $\sin^2(2\theta_{23}) = 1$  at 90 % C.L. (F&C)







# **Oscillation Analysis:** $u_{\mu} \rightarrow \nu_{e}$

Search for  $\nu_{\mu} \rightarrow \nu_{e}$  oscillations with the OPERA experiment in the CNGS beam, JHEP **1307** (2013) 004







### **ECC** reconstruction:



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### Systematic $\nu_e$ Event Selection



### CS em shower hints:





- Interpolation of 1ry vertex tracks to CS
- Expanded scan volume
- Analysis of downstream bricks

### Backgrounds:

- $\nu_e$  from intrinsic beam contamination
- $e^+e^-$  from  $\pi^0$  decays misidentified as single-e
- $u_{\tau}$  CC interactions with  $\tau \rightarrow e$

### $\nu_e$ Oscillation Analysis



### Final data sample (2008 – 2012): 1185 events

Contribution	Expected events $17.97 \times 10^{19}$ p.o.t.	
Beam contamination	30.8	
au  ightarrow e	0.9	
$\pi^0$	0.5	
$\nu_{\mu} \rightarrow \nu_{e}$	2.7	
Total	34.9	
Observed	34	
PRELIMINARY		



### **3+1 energy shape** analysis:



$$\sin^2 2\theta_{\mu e} = 4|U_{\mu 4}|^2|U_{e4}|^2$$



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## **Conclusion & Outlook**





### $u_{\mu} ightarrow u_{ au}$ oscillation analysis: Standard analysis

- 5  $\nu_{\tau}$  candidate events observed (0.25 BG events expected)
- $\triangleright$  Discovery of  $\nu_{\tau}$  appearance @ 5.1 $\sigma$

**Outlook:**  $\nu_{\mu} \rightarrow \nu_{\tau}$  **Minimum bias analysis** 

- Improved statistics and signal-to-noise ratio
- ▷ Measurement of  $\Delta m_{23}^2$  in appearance mode
- $\triangleright \quad \textbf{Measurement of } \nu_{\tau} \text{ cross section}$
- > Sterile neutrino analysis

**Outlook:**  $\nu_{\mu} \rightarrow \nu_{e}$  analysis (full data sample)

- Improved statistics and analysis method
- ▷ Sterile neutrino analysis: 3+1 energy shape analysis
- ▷ 3-flavour neutrino analysis



### **Conclusion & Outlook**



### Further studies:

- Combined analysis:  $\nu_{\mu} \rightarrow \nu_{\tau}, \nu_{\mu} \rightarrow \nu_{e}, \nu_{\mu} \rightarrow \nu_{\mu}$
- 0µ double decay event
- Charged particle multiplicity distributions
- Annual  $\mu$  rate modulation

### Future experiments:

- Improved nuclear emulsions & scanning techniques:
- $\triangleright~$  Muon radiography, directional DM search,  $\gamma$  telescopes...

### • SHiP:

- OPERA-like ECC bricks
- Drift tubes
- JUNO:
- OPERA RPC



### Thank you for your attention!



### 11 countries, 28 institutes, 140 physicists...

